

# Tables for finding the Latitude & the Direction of the Meridian.

## PRELIMINARY REMARKS AND COMPUTATIONS.

The mean astronomical day begins at mean noon, or twelve hours later than the ordinary civil day; and mean astronomical time is expressed by the number of hours, minutes, and seconds that have elapsed since the preceding mean noon.

given, it is required to find the date in astronomical mean time.

*Rule.*—If the civil time be P. M., the astronomical time is the same as the civil time, (omitting the sign P. M.)

If the civil time be A. M., twelve hours must be added to the hours, and the date must be put one day back.

*Prob. 1.*—A date in local civil time being

*Examples:*— DATES IN CIVIL TIME. CORRESPONDING ASTRONOMICAL TIME.

	h. m.		h. m.
April 7,	5 ½	P. M.	April 7,
May 5,	10 17	A. M.	May 4,
July 1,	3 40	A. M.	June 30,
			15 40

*Prob. 2.*—The longitude being given in degrees, &c., it is required to express the same in time, at the rate of 1 hour for 15 degrees, &c.

seconds by 15, or, what is more convenient, multiply by 4 and divide the product by 60.

*Example.*—The longitude being 80° 20' 36", find the longitude in time.

*Rule.*—Divide the degrees, minutes, and

$$\begin{array}{r}
 80^{\circ} \ 20' \ 36'' \\
 \underline{\hspace{1.5cm}} \\
 60 \ ) \ 321 \ 22 \ 24 \\
 \underline{\hspace{1.5cm}} \\
 5 \ 21 \ 22.4
 \end{array}$$

h. m. s.  
Long. in time is 5 21 22.4

*Prob. 3.*—The date being given in local time, (civil or astronomical,) required to find the Greenwich date, the longitude being west.

the day being unaltered; but if the sum be more than 24 hours, 24 hours must be rejected, and the date must be put one day forward.

*Rule.*—To the local time expressed astronomically add the long. in time. The sum, if less than 24 hours, will be the Greenwich date,

*Examples.*—Required the Greenwich dates corresponding to the following local times.

	h. m.		° ' "
(1) Dec. 20	6 14	P. M.	Long. 42 15 W.
(2) March 12	11 20	A. M.	Long. 80 30 W.
(1)			(2)
Place Dec 20	6 14		Place March 11
Long. in time	2 49 W.		Long. in time
	<hr style="width: 50%; margin: 0 auto;"/>		h. m.
Gr. Dec. 20	9 3		23 20
	<hr style="width: 50%; margin: 0 auto;"/>		5 22
			<hr style="width: 50%; margin: 0 auto;"/>
		Gr. March 11	28 42
		or Gr. March 12	4 42

*Prob. 4.*—To find from table I the declination of the sun at apparent noon, i.e. at the instant when the sun passes the meridian of the place, the long. being west.

meridian of Greenwich to the meridian of the place.

*Rule.*—(1) Take from table I the Declination at Greenwich apparent noon of the proposed day, and the hourly change of Declination from the column marked (D), giving to each their proper signs N. or S.

(3) If the declination and its change have the same sign (N or S), add them together and give to the sum their common sign; if their signs be contrary, subtract the less from the greater and give to the remainder the sign of the greater. The sum or difference, as the case may be, will be the required declination at local apparent noon.

(2) Express the longitude in time, and multiply (D) by the hours and parts of an hour. This will be the change which the declination undergoes while the sun passes from the me-

*Examples.*—Find the declination of the sun at apparent noon in the following cases:

	h. m.		° ' "
(1) April 10, 1872		Long. 80 45;	
(2) January 5, "		" 70 15;	
(1)			(2)
Long. in time	5 23		h. m.
	<hr style="width: 50%; margin: 0 auto;"/>		Long. in time 4 41
April 10 Declin. Gr. app. noon	8 10	12.3 N	Jan. 5, 22 39 49.3 S
Change in 1 h.	55.24	N	16.55 N
" 5 23	4 57.4	N	1 17.5 N
	<hr style="width: 50%; margin: 0 auto;"/>		<hr style="width: 50%; margin: 0 auto;"/>
Decl. local app. noon	8 15 9.7	N	22 38 31.8 S

Tables II and III are designed to aid in finding the time when a given bright star passes the meridian on any proposed night; and also in finding what stars pass the meridian between two proposed hours.

Table II contains sixty-nine stars of magnitudes not less than the third, and placed in the order of their culmination. Polaris and  $\beta$  Ursæ Minoris, as they occur in another table,

are not included in table II. For convenience of reference the stars are distinguished by the Roman numerals. The Right Ascensions (R. A.) and the declinations, refer to their mean places on Jan. 1, 1872, at midnight. Although not accurately true, they are sufficiently near the truth for ordinary purposes. Where very great exactness is needed the Nautical Almanac must be employed.